Generating Method Based on Conformal Invariance of the Maxwell Field

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ERE 2010







Motivation

- Generating Technique
- Existing Theorems
- Our Generating Method
 - Introducing the Method
 - Application



Generating Technique Existing Theorems

Outline



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Generating Technique Existing Theorems

Generating Methods in General

Exact solutions are difficult to find

Generating method:

- Take an existing solution (seed spacetime)
- Apply some transformation
- Obtain a new solution (*new spacetime*)
- Method we use: conformal transformation



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Generating Technique Existing Theorems

Conformal Transformation

• Seed spacetime: $g_{\mu\nu}$

- 2 Multiply metric by suitable scalar function Ω^2
- 3 New spacetime: $\widetilde{g}_{\mu
 u}=\Omega^2 g_{\mu
 u}$

Note: Maxwell field is conformally invariant!

• If $F_{\mu\nu}$ is a solution of source-free Maxwell equations on background $g_{\mu\nu}$, it is also a solution on $\tilde{g}_{\mu\nu}$



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- 2 Multiply metric by suitable scalar function Ω^2
- (a) New spacetime: $\tilde{g}_{\mu\nu} = \Omega^2 g_{\mu\nu}$

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Theorems Concerning Conformal Transformation

Brinkmann

The only properly conformally related Einstein spaces $(R_{\mu\nu} = \Lambda g_{\mu\nu})$ are vacuum pp-waves or Minkowsky and (A)dS.

Daftardar-Gejji

If $\tilde{G}_{\mu\nu} = G_{\mu\nu}$ then both spacetimes are pp-waves.

Many other theorems suggesting that vacuum is not very suitable seed spacetime...



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Introducing the Method Application

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Introducing the Method Application

Main Idea

- Seed spacetime: Solution of Einstein-Maxwell equations $(g_{\mu\nu}, F_{\mu\nu})$
- ② Conformal transformation $ilde{g}_{\mu
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- New spacetime: Solution of Einstein-Maxwell equations (*ğ*_{μν}, *F*_{μν})
- Maxwell equations are automatically fulfilled
- Einstein equations impose conditions on Ω
- Existing theorems do not apply here



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Introducing the Method

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Introducing the Method

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Introducing the Method Application

Einstein equations

• Einstein equations in the new spacetime $(\tilde{g}_{\mu\nu}, F_{\mu\nu})$:

$$ilde{R}_{\mu
u} = \mathbf{8}\pi\, ilde{T}_{\mu
u}$$

• Transformation properties:

$$\begin{split} \tilde{R}_{\mu\nu} &= R_{\mu\nu} - \frac{2}{\Omega} \Omega_{;\mu\nu} + \frac{4}{\Omega^2} \Omega_{,\mu} \Omega_{,\nu} - \frac{1}{\Omega} \Box \Omega g_{\mu\nu} - \frac{1}{\Omega^2} \Omega_{,\rho} \Omega^{;\rho} g_{\mu\nu} \\ \tilde{T}_{\mu\nu} &= \Omega^{-2} T_{\mu\nu} \end{split}$$

Introducing the Method Application

Conditions on the conformal factor

• Resulting equations:

$$(\Omega^2 - 1)R_{\mu\nu} - 2\Omega\Omega_{;\mu\nu} + 4\Omega_{,\mu}\Omega_{,\nu} - g_{\mu\nu}\Omega_{,\rho}\Omega^{;\rho} = 0 \qquad (1)$$

- Trace $\Rightarrow \Box \Omega = 0$
- Overdetermined system: 10 equations for one function Ω , generally no solution (except for $\Omega^2 \equiv 1$)

Introducing the Method Application

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Application



Introducing the Method Application

Seed Spacetime

- Most seed spacetimes do not admit non-trivial Ω
- The only suitable family of spacetimes we found so far: pp-waves

$$ds^2 = -2H(u,\xi,ar{\xi})du^2 - 2dudv + 2d\xi dar{\xi}$$

Note: Vector field $\mathbf{k} := \partial / \partial v$ is covariantly constant



Introducing the Method Application

Conformal Factor

- Eqs. (1) have a non-trivial solution iff H(u, ξ, ξ) = h(u)ξξ in suitable coordinates and Ω = Ω(u)
- Such pp-wave is conformally flat
- Eqs. (1) then reduce to an ODE:

$$\Omega \frac{d^2 \Omega}{du^2} - 2 \left(\frac{d\Omega}{du} \right)^2 + h(u)(1 - \Omega^2) = 0$$
 (2)



Introducing the Method Application

New Spacetime

- k := ∂/∂v remains covariantly constant ⇒ new spacetime is also a conformally flat pp-wave
- By construction: both *pp*-waves have the same $F_{\mu\nu}$
- Question: Is the new spacetime different from the seed?



Introducing the Method Application

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Introducing the Method Application

(Non)Equivalence

• Explicit choice of seed

$$ds^2=-4\xiar\xi du^2-2dudv+2d\xi dar\xi$$

Note:
$$R_{\mu
u} = 4k_{\mu}k_{
u} \Rightarrow R_{\mu
u;
ho} = 0$$

• Explicit solution of eq. (2): $\Omega = tanh(u)$, i.e.

$$\widetilde{ds}^2 = \tanh^2(u)(-4\xi\overline{\xi}du^2 - 2dudv + 2d\xi d\overline{\xi})$$

Note
$$\tilde{R}_{\mu\nu} = 4 \coth^2(u) k_{\mu} k_{\nu} \Rightarrow \tilde{R}_{\mu\nu;\rho} \neq 0$$

Answer: Spacetimes are not equivalent!

Introducing the Method Application

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- We presented a generating method that produces a solution to Einstein-Maxwell equations if a suitable seed spacetime is found
- Once again, *pp*-waves proved to play an important role in context of conformal transformation
- Outlook
 - Are there any other suitable seeds other than pp-waves?
 - Can the method be generalized to be less restrictive?