Entropy of an Evolving Black Hole:



Clues from hydrodynamics

ERE 2010 - Granada

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PRD 80, 126013 (2009); <u>arXiv:0910.0748</u>

(Some figures and ideas from: I.B. and Jonathan Martin, <u>arXiv:1007.1642</u>)

Black holes in equilibrium

area of horizon



 Is entropy still well defined (at least in some quasi-equilibrium regime)?

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 If so is it still proportional to horizon area?

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• If so which horizon?

Difficulties with event horizons



 Locally they are null hypersurfaces ruled by null geodesics

$$\mathcal{L}_{\ell}\theta_{(\ell)} = \kappa_{\ell}\theta_{(\ell)} - \|\sigma^{(\ell)}\|^2$$
$$-G_{ab}\ell^a\ell^b - \frac{\theta_{(\ell)}^2}{(n-1)}$$

 But exactly which hypersurface is defined by future boundary conditions

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Difficulties with apparent horizons



 Hypersurface foliated with surfaces of vanishing outward null expansion:

$$\theta_{(\ell)} = \tilde{q}^{ab} \nabla_a \ell_b = 0$$

 Can be (quasi)locally identified. Local dynamics.

$$\mathcal{L}_{\mathcal{V}}\theta_{(\ell)} = \mathcal{L}_{\ell}\theta_{(\ell)} + \kappa_{\mathcal{V}}\theta_{(\ell)} - d^2C + 2\tilde{\omega}^a d_aC$$
$$-C\left[||\tilde{\omega}||^2 - d_a\tilde{\omega}^a - \tilde{R}/2 + G_{ab}\ell^a n^b - \theta_{(\ell)}\theta_{(n)}\right]$$

• Are NOT unique.

Difficulties with apparent horizons



 Hypersurface foliated with surfaces of vanishing outward null expansion:

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 Can be (quasi)locally identified. Local dynamics.

$$0 = \mathcal{L}_{\ell} \theta_{(\ell)} - d^2 C + 2 \tilde{\omega}^a d_a C$$
$$C \left[||\tilde{\omega}||^2 - d_a \tilde{\omega}^a - \tilde{R}/2 + G_{ab} \ell^a n^b \right]$$

• Are NOT unique.

Difficulties with apparent horizons



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apparent

 Σ_t

Quasi-equilbrium regime



 Characterized by slowly evolving apparent horizons (BF 03,07)

$$\epsilon = \frac{\mathcal{L}_{\hat{\mathcal{V}}}\sqrt{\tilde{q}}}{\sqrt{\tilde{q}}} = -\sqrt{\frac{C}{2}}\theta_{(n)} \ll 1$$

• First law holds (probably all horizons)

$$\kappa_o \mathcal{L}_{\mathcal{V}} \sqrt{\tilde{q}} \approx \sqrt{\tilde{q}} \left(||\sigma^{(\ell)}||^2 + G_{ab} \ell^a \ell^b \right)$$

- Various horizons are close (BM 10, Nielsen 10)
- Most black hole physics is in this regime including some black hole formations (BK 05)

AdS-CFT – The hydrodynamic regime

• In the long-wavelength limit the CFT becomes a theory of conformally invariant fluid mechanics:



$$abla_{\mu}T^{\mu\nu}_{\mathscr{I}} = 0 \quad \text{and} \quad \nabla_{\mu}J^{\mu}_{I} = 0$$

+ invariance under $g_{\mu\nu} \rightarrow e^{2\phi}\tilde{g}_{\mu\nu}$

- Fluid flows on boundary are dual to black brane/hole solutions
- Can be derived from a perturbed GR solution, independent of the original conjecture.(Bhattacharya, Hubeny, Minwalla, Rangamani 07)
- Quasi-equilibrium regime (on both sides)
- Thermodynamics should match

BUT.....

• The boundary is dual to the FULL spacetime.



- How do you match bulk to boundary?
- Thermodynamics isn't uniquely defined on the boundary either...
 - thermodynamic quantities must be consistent with the symmetries
 - entropy flow is a vector field with non-negative divergence

 $\nabla_{\mu}S^{\mu} \ge 0$

• do the uncertainties match?

Bjorken flow

• A particularly simple example is boost-invariant flow:



4D flow with boost invariance in direction of motion and planar symmetry perpendicular to motion

• All variables depend on proper time alone (large τ)

$$\varepsilon = e_o T^4$$
, $s = \frac{4e_o}{3}T^3$ and $T \approx \frac{\Lambda}{\tau^{1/3}}$

 Dual is a (perturbed) boosted black brane

$$ds^{2} = -r^{2}A(\tau, r)d\tau^{2} + 2d\tau dr + (1+\tau r)^{2}e^{b(\tau, r)}dy^{2} + r^{2}e^{c(\tau, r)}dx_{\perp}^{2}$$



Quark-gluon plasma at RHIC (2000)

Bjorken flow – gravity vs fluid

• For apparent horizons, the formalisms nicely match:

i) natural expansion parameter is $\epsilon = rac{1}{ au^{2/3}}$

ii) surface gravity is $\kappa = \frac{2\pi\Lambda}{\tau^{1/3}}$

iii) first law: $\kappa_{\mathcal{V}} \mathcal{L}_{\mathcal{V}} \sqrt{\tilde{q}} = \sqrt{\tilde{q}} ||\sigma_{(\ell)}||^2 \approx \frac{2\pi^2 \Lambda^3}{3} \cdot \frac{1}{\tau^2}$

• Entropy:

$$\begin{array}{ll} \textbf{i) fluid:} & s_{bnd} \propto \frac{\Lambda^3}{\tau} \left\{ 1 - \frac{1}{2\pi\Lambda} \cdot \frac{1}{\tau^{2/3}} + \frac{0.008\beta}{\Lambda^2} \cdot \frac{1}{\tau^{4/3}} + O\left(\frac{1}{\tau^2}\right) \right\} \\ \textbf{ii) AH:} & s_{AH} = \frac{N_c^2 \pi^2}{2} \cdot \frac{\Lambda^3}{\tau} \left\{ 1 - \frac{1}{2\pi\Lambda} \cdot \frac{1}{\tau^{2/3}} + \frac{0.039}{\Lambda^2} \cdot \frac{1}{\tau^{4/3}} + O\left(\frac{1}{\tau^2}\right) \right\} \\ \textbf{iii) EH:} & s_{EH} = \frac{N_c^2 \pi^2}{2} \cdot \frac{\Lambda^3}{\tau} \left\{ 1 - \frac{1}{2\pi\Lambda} \cdot \frac{1}{\tau^{2/3}} + \frac{0.056}{\Lambda^2} \cdot \frac{1}{\tau^{4/3}} + O\left(\frac{1}{\tau^2}\right) \right\} \end{array}$$

Why the discrepancy?

- Answer #1: (practical entropy) Either horizon is fine. For quasiequilibrium thermodynamics, don't worry about second order. The horizons are arbitrarily close (BM 10).
- Answer #2: (hydro dynamic unknowns) There is some (unknown) way to fix the fluid entropy flow. Then it would match that of one horizon or the other.
- Answer #3: (bulk corrections) There are other corrections needed in the bulk...
- Answer #3: (phenomenological entropy) Ambiguity on the boundary matches an ambiguity in the bulk. One possibility is to associate entropy with "almost" null, "almost" vanishing $\theta_{(\ell)}$, expanding surfaces (with correct limits).

Conclusions

- There is an excellent match between the bulk gravity and boundary fluid thermodynamics.
- The uncertainty on the boundary does NOT match the ambiguity in AH location.
- That uncertainty may suggest a rethinking of entropy in the bulk.
- A more general calculation is on its way...